

# Aspects of Diffeomorphism Invariant Theory of Extended Objects

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The structure of a diffeomorphism invariant Lagrangians for an extended object  $W$  embedded in a bulk space  $M$  is discussed by following a close analogy with the relativistic particle in electromagnetic field as a system that is reparametrization-invariant. The current construction naturally contains, relativistic point particle, string theory, and Dirac–Nambu–Goto Lagrangians with Wess–Zumino terms. For Lorentzian metric field, the non-relativistic theory of an integrally submerged  $W$ -brane is well defined provided that the brane does not alter the background interaction fields. A natural time gauge is fixed by the integral submergence (sub-manifold structure) within a Lorentzian signature structure. A generally covariant relativistic theory for the discussed brane Lagrangians is also discussed. The mass-shell constraint and the Klein–Gordon equation are shown to be universal when gravity-like interaction is present. A construction of the Dirac equation for the  $W$ -brane that circumvents some of the problems associated with diffeomorphism invariance of such Lagrangians by promoting the velocity coordinates into a non-commuting gamma variables is presented.

**Keywords:** diffeomorphism invariant systems, reparametrization-invariant systems, matter Lagrangian, homogeneous singular Lagrangians, relativistic particle, Dirac equation, string theory, extended objects, branes, interaction fields, generally covariant theory, gauge symmetries, background free theories.

**Introduction.** The Hamiltonian and Lagrangian formulation[1, 2] are two very useful approaches in physics. In general, these two approaches are related by the Legander transformation. For a reparametrization-invariant theory, however, there are problems in changing from the Lagrangian to the Hamiltonian approach.[2, 3, 4, 5] In this paper the focus is on the properties of reparametrization-invariant matter systems such as, the relativistic particle and its extended object (brane) generalization within the Lagrangian approach. We try to answer the question: “What is the Lagrangian for an extended ‘matter’ object?”

**Matter Lagrangian for relativistic particle.** The action for a massive relativistic particle has a nice geometrical meaning: it is the distance along the particle trajectory[6] provided that the units are such that  $x^0 = ct$  and the particle moves with a constant 4-velocity ( $g_{\mu\nu}v^\mu v^\nu = 1$ ):

$$S_1 = \int d\tau L_1(x, v) = \int d\tau \sqrt{g_{\mu\nu}v^\mu v^\nu} \rightarrow \int d\tau. \quad (1)$$

For a massless particle, such as a photon, the length of the 4-velocity is zero ( $g_{\mu\nu}v^\mu v^\nu = 0$ ) and the appropriate ‘good’ action[6] is:

$$S_2 = \int L_2(x, v)d\tau = \int g_{\mu\nu}v^\mu v^\nu d\tau. \quad (2)$$

The Euler–Lagrange equations obtained from  $S_1$  and  $S_2$  are equivalent, even more, they are equivalent to the geodesic equation as well:

$$\frac{d}{d\tau} \vec{v} = D_{\vec{v}} \vec{v} = v^\beta \nabla_\beta \vec{v} = 0 \quad (3)$$

Since the Levi–Civita connection  $\nabla$  preserves the length of the vectors[6] ( $\nabla g(\vec{v}, \vec{v}) = 0$ ) this equivalence is not surprising because the Lagrangians in (1) and (2) are functions of the preserved arc length  $g(\vec{v}, \vec{v}) = \vec{v}^2$ . However, the equivalence between  $S_1$  and  $S_2$  has a much deeper roots.

**Homogeneous Lagrangians.** Since  $L_2$  is a homogeneous function of order 2 with respect to  $\vec{v}$ , the corresponding Hamiltonian function ( $h = v^\beta \partial L / \partial v^\beta - L$ ) is exactly equal to  $L_2$ . Thus  $L_2$  is conserved, and so is the length of  $\vec{v}$ . Any

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homogeneous Lagrangian in  $\vec{v}$  of order  $n \neq 1$  is conserved because  $h = (n-1)L$ . If  $dL/d\tau = 0$ , then the Euler–Lagrange equations for  $L$  and  $\tilde{L} = f(L)$  are equivalent. This is an equivalence that applies to homogeneous Lagrangians in particular. It is different from the usual equivalence  $L \rightarrow \tilde{L} = L + d\Lambda/d\tau$  or the more general equivalence discussed by Hojman and Harleston Ref. [7]. Any solution of the Euler–Lagrange equation for  $\tilde{L} = L^\alpha$  would conserve  $L = L_1$  since  $\tilde{h} = (\alpha - 1)L^\alpha$ . All these solutions are solutions of the Euler–Lagrange equation for  $L$  as well, thus  $L^\alpha \subset L$ . In general, conservation of  $L_1$  is not guaranteed since  $L_1 \rightarrow L_1 + d\Lambda/d\tau$  is also a homogeneous Lagrangian of order one equivalent to  $L_1$ . This suggests that there may be a choice of  $\Lambda$ , a “gauge fixing”, such that  $L_1 + d\Lambda/d\tau$  is conserved even if  $L_1$  is not. The above discussion applies to a more general homogeneous Lagrangians as well.[8]

In the example of the relativistic particle, the Lagrangian and the trajectory parameterization have a geometrical meaning. In general, however, parameterization of a trajectory is quite arbitrary for any observer. If there is no preferred trajectory parameterization in a smooth space-time, then we are free to choose the standard of distance (time, using natural units  $c = 1$ ). Thus, *our theory should not depend on the choice of parameterization*. By inspection of the Euler–Lagrange equations, any homogeneous Lagrangian of order  $n$  ( $L(x, \alpha\vec{v}) = \alpha^n L(x, \vec{v})$ ) provides a reparametrization invariant equations ( $\tau \rightarrow \tau/\alpha, \vec{v} \rightarrow \alpha\vec{v}$ ). Next, note that the action  $S$  involves an integration that is a natural structure for orientable manifolds ( $M$ ) with an  $n$ -form of the volume. Since a trajectory is a one-dimensional object, then what we are looking at is an embedding  $\phi : \mathbb{R}^1 \rightarrow M$ . This means that we push forward the tangential space  $\phi_* : T(\mathbb{R}^1) = \mathbb{R}^1 \rightarrow T(M)$ , and pull back the cotangent space  $\phi^* : T(\mathbb{R}^1) = \mathbb{R}^1 \leftarrow T^*(M)$ . Thus a 1-form  $\omega$  on  $M$  that is in  $T^*(M)$  ( $\omega = A_\mu(x) dx^\mu$ ) will be pulled back on  $\mathbb{R}^1$  ( $\phi^*(\omega)$ ) and there it should be proportional to the volume form on  $\mathbb{R}^1$  ( $\phi^*(\omega) = A_\mu(x) (dx^\mu/d\tau) d\tau \sim d\tau$ ), allowing us to integrate  $\int \phi^*(\omega) :$

$$\int \phi^*(\omega) = \int L d\tau = \int A_\mu(x) v^\mu d\tau.$$

Therefore, by selecting a 1-form  $\omega = A_\mu(x) dx^\mu$  on  $M$  and using  $L = A_\mu(x) v^\mu$  we are actually solving for the embedding  $\phi : \mathbb{R}^1 \rightarrow M$  using a chart on  $M$  with coordinates  $x : M \rightarrow \mathbb{R}^n$ . The Lagrangian obtained this way is homogeneous of first order in  $v$  with a very simple dynamics. The corresponding Euler–Lagrange equation is  $F_{\nu\mu} v^\mu = 0$  where  $F$  is a 2-form ( $F = dA$ ) – the Faraday’s tensor. If the assumption that  $L$  is a pulled back 1-form is relaxed and instead one assumes that it is just a homogeneous Lagrangian of order one, then the corresponding reparametrization-invariant theory may have an interesting dynamics.

**First order homogeneous Lagrangians – canonical form.** Now we define what we mean by the *canonical form of the first order homogeneous Lagrangian* and why do we prefer this mathematical expression. Let  $S_{\alpha_1\alpha_2\dots\alpha_n}$  be a symmetric tensor of rank  $n$  which defines a homogeneous function of order  $n$  ( $S_n(\vec{v}, \dots, \vec{v}) = S_{\alpha_1\alpha_2\dots\alpha_n} v^{\alpha_1} \dots v^{\alpha_n}$ ). The symmetric tensor of rank two is denoted by  $g_{\alpha\beta}$ . Using this notation, the canonical form of the first order homogeneous Lagrangian is defined as:

$$L(\vec{x}, \vec{v}) = \sum_{n=1}^{\infty} \sqrt[n]{S_n(\vec{v}, \dots, \vec{v})} = A_\alpha v^\alpha + \sqrt{g_{\alpha\beta} v^\alpha v^\beta} + \dots \sqrt[n]{S_n(\vec{v}, \dots, \vec{v})}. \quad (4)$$

Any Lagrangian for the matter should involve interaction fields that couple with the velocity  $\vec{v}$  to a scalar. When the matter action is combined with the action for the interaction fields ( $\mathcal{S} = \int \mathcal{L} dV$ ), one obtains a full *background independent theory*. Then the corresponding Euler–Lagrange equations contain “dynamical derivatives” on the left hand side and sources on the right hand side:

$$\partial_\gamma \left( \frac{\delta \mathcal{L}}{\delta(\partial_\gamma \Psi^\alpha)} \right) = \frac{\delta \mathcal{L}}{\delta \Psi^\alpha} + \frac{\partial L_{\text{matter}}}{\partial \Psi^\alpha}.$$

The advantage of the canonical form of the first order homogeneous Lagrangian (4) is that each interaction field, which is associated with a symmetric tensor, has a unique matter source that is a monomial in the velocities:

$$\frac{\partial L}{\partial S_{\alpha_1\alpha_2\dots\alpha_n}} = \frac{1}{n} (S_n(\vec{v}, \dots, \vec{v}))^{\frac{1-n}{n}} v^{\alpha_1} \dots v^{\alpha_n}. \quad (5)$$

There are many other ways one can write first-order homogeneous functions.[4] For example, one can consider the following expression  $L(\vec{x}, \vec{v}) = (h_{\alpha\beta} v^\alpha v^\beta) (g_{\alpha\beta} v^\alpha v^\beta)^{-1/2}$ . However, each of the fields  $h$  and  $g$  has the same source type ( $\sim v^\alpha v^\beta$ ). At this stage, however, it is not clear why the same source type should produce different fields. Therefore, the canonical form (4) seems more appropriate for the current discussion.

**Extended objects.** In the previous sections, the classical mechanics of a point-like particle have been discussed as a problem concerned with the embedding  $\phi : \mathbb{R}^1 \rightarrow M$ . The map  $\phi$  provides the trajectory (the word line) of the

particle in the target space  $M$ . In this sense, we are dealing with a one dimensional object, the world-line of the particle (one dimensional W-brane). We think of an extended object as a manifold  $W$  with dimension denoted by  $D$ . In this sense, we have to solve for  $\phi : W \rightarrow M$  such that some action integral is minimized. From this point of view, we are dealing with mechanics of a brane. In other words, how is this  $D$ -dimensional extended object submerged in  $M$ , and what are the relevant interaction fields? Following the relativistic point particle discussion, we consider the space of the  $D$ -forms over the manifold  $M$ , denoted by  $\Lambda^D(M)$ , that has dimension  $\binom{m}{D} = \frac{m!}{D!(m-D)!}$ . An element  $\Omega$  in  $\Lambda^D(M)$  has the form  $\Omega = \Omega_{\alpha_1 \dots \alpha_m} dx^{\alpha_1} \wedge dx^{\alpha_2} \wedge \dots \wedge dx^{\alpha_m}$ . We use the label  $\Gamma$  to index different  $D$ -forms over  $M$ ,  $\Gamma = 1, 2, \dots, \binom{m}{D}$ ; thus  $\Omega \rightarrow \Omega^\Gamma = \Omega_{\alpha_1 \dots \alpha_m}^\Gamma dx^{\alpha_1} \wedge dx^{\alpha_2} \wedge \dots \wedge dx^{\alpha_m}$ . Next we introduce “generalized velocity vectors” with components  $\omega^\Gamma$ :

$$\omega^\Gamma = \frac{\Omega^\Gamma}{dz} = \Omega_{\alpha_1 \dots \alpha_D}^\Gamma \frac{\partial(x^{\alpha_1} x^{\alpha_2} \dots x^{\alpha_D})}{\partial(z^1 z^2 \dots z^D)}, \quad dz = dz^1 \wedge dz^2 \wedge \dots \wedge dz^D.$$

In the above expression,  $\frac{\partial(x^{\alpha_1} x^{\alpha_2} \dots x^{\alpha_D})}{\partial(z^1 z^2 \dots z^D)}$  represents the Jacobian of the transformation from coordinates  $\{x^\alpha\}$  over the manifold  $M$  to coordinates  $\{z^a\}$  over the brane. The pull back of a  $D$ -form  $\Omega^\Gamma$  must be proportional to the volume form over the brane:

$$\phi^*(\Omega^\Gamma) = \omega^\Gamma dz^1 \wedge \dots \wedge dz^D = \Omega_{\alpha_1 \dots \alpha_D}^\Gamma \frac{\partial(x^{\alpha_1} \dots x^{\alpha_D})}{\partial(z^1 \dots z^D)} dz^1 \wedge \dots \wedge dz^D.$$

Thus, it is suitable for integration over the  $W$ -manifold and the action is:

$$S[\phi] = \int_W L(\vec{\phi}, \vec{\omega}) dz = \int_W \phi^*(\Omega) = \int_W A_\Gamma(\vec{\phi}) \omega^\Gamma dz.$$

This is a homogeneous function in  $\omega$  and is reparametrization (diffeomorphism) invariant with respect to the diffeomorphisms of the  $W$ -manifold. If we relax the linearity  $L(\vec{\phi}, \vec{\omega}) = \phi^*(\Omega) = A_\Gamma(\vec{\phi}) \omega^\Gamma$  in  $\vec{\omega}$ , then the canonical expression for the first order homogeneous Lagrangian is:

$$L(\vec{\phi}, \vec{\omega}) = \sum_{n=1}^{\infty} \sqrt[n]{S_n(\vec{\omega}, \dots, \vec{\omega})} = A_\Gamma \omega^\Gamma + \sqrt{g_{\Gamma_1 \Gamma_2} \omega^{\Gamma_1} \omega^{\Gamma_2}} + \dots \quad (6)$$

At this point, there is a strong analogy between the relativistic point particle and the extended object. Some specific examples of  $W$ -brane theories correspond to the following familiar Lagrangians:

*Lagrangian for a relativistic point particle* in an electromagnetic field:  $\dim W = 1$  (World-line) with  $\omega^\Gamma \rightarrow v^\alpha = \frac{dx^\alpha}{d\tau}$

$$L(\vec{\phi}, \vec{\omega}) = A_\Gamma \omega^\Gamma + \sqrt{g_{\Gamma_1 \Gamma_2} \omega^{\Gamma_1} \omega^{\Gamma_2}} \rightarrow q A_\alpha v^\alpha + m \sqrt{g_{\alpha\beta} v^\alpha v^\beta}.$$

*Lagrangian for strings:*  $\dim W = 2$  (World-sheet)

$L(x^\alpha, \partial_\alpha x^\beta) = \sqrt{Y^{\alpha\beta} Y_{\alpha\beta}}$ , with the following notation:

$$\omega^\Gamma \rightarrow Y^{\alpha\beta} = \frac{\partial(x^\alpha, x^\beta)}{\partial(\tau, \sigma)} = \det \begin{pmatrix} \partial_\tau x^\alpha & \partial_\sigma x^\alpha \\ \partial_\tau x^\beta & \partial_\sigma x^\beta \end{pmatrix} = \partial_\tau x^\alpha \partial_\sigma x^\beta - \partial_\sigma x^\alpha \partial_\tau x^\beta.$$

*Dirac–Nambu–Goto Lagrangian (DNG)* [9]:  $L(x^\alpha, \partial_W x^\beta) = \sqrt{Y^\Gamma Y_\Gamma}$ .

The corresponding electromagnetic interaction term for W-branes is known as Wess–Zumino term [10] in string theory.

From the expressions (4) and (6), one can see that the corresponding matter Lagrangians ( $L$ ), in their canonical form, contain electromagnetic ( $A$ ) and gravitational ( $g$ ) interactions, as well as interactions that are not clearly identified yet ( $S_n$ ,  $n > 2$ ), if present at all in nature. At this stage, we have a theory with background fields since we don't know the equations for the interaction fields  $A$ ,  $g$ , and  $S_n$ . To complete the theory, we need to introduce actions for these interaction fields. If one is going to study the new interaction fields  $S_n$ ,  $n > 2$ , then some guiding principles for writing field Lagrangians are needed.

One approach is to apply the above discussion and view the  $S_n$  fields as related to an embedding of the  $M$ -manifold into the manifold of symmetric tensors over  $M$ . Another approach would be to use the external derivative  $d$ , external multiplication  $\wedge$ , and Hodge dual  $*$  operations in the external algebra  $\Lambda(T^*M)$  over  $M$  to construct objects

proportional to the volume form over  $M$ . For example, for any  $n$ -form ( $A$ ) the expressions  $A \wedge *A$  and  $dA \wedge *dA$  are forms proportional to the volume form. The next important principle comes from the symmetry in the matter equation. That is, if there is a transformation  $A \rightarrow A'$  that leaves the matter equations unchanged, then there is no way to distinguish  $A$  and  $A'$ . Thus the action for the field  $A$  should obey the same gauge symmetry. For the electromagnetic field ( $A \rightarrow A' = A + df$ ) this leads uniquely to the field Lagrangian  $\mathcal{L} = dA \wedge *dA = F \wedge *F$ , when for gravity[11] it leads to the Cartan–Einstein action[12]  $S[R] = \int R_{\alpha\beta} \wedge *(*dx^\alpha \wedge dx^\beta)$ .

**Non-relativistic limit.** For a  $W$ -brane we assume the existence a local coordinate frame where one component of the generalized velocity can be set to 1 ( $\omega^0 = 1$ ). This generalized velocity component is associated with the brane “time coordinate.” In fact,  $\omega^0 = 1$  means that there is an integral embedding of the brane in the target space  $M$ , and the image of the brane is a sub-manifold of  $M$ . If the coordinates of  $M$  are labeled so that  $x^i = z^i, i = 1, \dots, D$ , then  $x^i$  are internal coordinates that can collapse to only one coordinate – the “world line”. This provides a gauge-fixing that allows one to do canonical quantization. This approach is mainly concerned with the choice of a coordinate time that is used as the trajectory parameter.[13, 14, 15, 16] Such choice removes the reparametrization invariance of the theory.

In a local coordinate system where  $\omega^0 = 1$  and the metric is a “one-time-metric” we have:

$$\begin{aligned} L &= A_\Gamma \omega^\Gamma + \sqrt{g_{\Gamma_1 \Gamma_2} \omega^{\Gamma_1} \omega^{\Gamma_2}} + \dots + \sqrt[m]{S_m(\vec{\omega}, \dots, \vec{\omega})} \rightarrow \\ &\rightarrow A_0 + A_i \omega^i + \sqrt{1 - g_{ii} \omega^i \omega^i} + \dots \approx A_0 + A_i \omega^i + 1 - \frac{1}{2} g_{ii} \omega^i \omega^i + \dots. \end{aligned}$$

Thus the Hamiltonian function is not zero anymore, so we can do canonical quantization, and the Hilbert space consists of the functions  $\Psi(x) \rightarrow \Psi(z, \tilde{x})$  where  $\tilde{x} = x^i, i = D+1, \dots, m$ . The brane coordinates  $z$  should be treated as  $t$  in quantum mechanics in the sense that the scalar product should be an integral over the space coordinates  $\tilde{x}$ . For  $W$ -branes the one-time coordinate reflects separation of the internal from the external coordinates when the  $W$ -brane is considered as a sub-manifold of the target space manifold  $M$ .

Even though canonical quantization can be applied after the above gauge fixing, one is not usually happy because the covariance of the theory is lost and time is a privileged coordinate. In general, there are well developed procedures for covariant quantization.[13, 15, 17, 18, 19] In this paper, however, we are not going to discuss these methods. Instead, we will employ a different quantization strategy[20], but before that we will discuss the mass-shell and Klein–Gordon equations.

**The mass-shell and Klein–Gordon equation.** Since the functional form of the canonical Lagrangian is the same for any  $W$ -brane, we use  $v$ , but it could be  $\omega$  as well. We define the momentum  $p$  and generalized momentum  $\pi$  for our canonical Lagrangian as follow:

$$\begin{aligned} p_\Gamma &= \frac{\delta L(\phi, \omega)}{\delta \omega^\Gamma} = e A_\Gamma + m \frac{g_{\Gamma \Sigma} \omega^\Sigma}{\sqrt{g(\vec{\omega}, \vec{\omega})}} + \dots + \frac{S_{\Gamma \Sigma_1 \dots \Sigma_n} \omega^{\Sigma_1} \dots \omega^{\Sigma_n}}{(S(\omega, \dots, \omega))^{1-1/n}} + \dots, \\ \pi_\alpha &= p_\alpha - e A_\alpha - \dots \frac{S_{\alpha \beta_1 \dots \beta_n} v^{\beta_1} \dots v^{\beta_n}}{(S(v, \dots, v))^{n/(n+1)}} \dots = m \frac{g_{\alpha \beta} v^\beta}{\sqrt{g(\vec{v}, \vec{v})}}. \end{aligned}$$

In the second equation we have used  $v$  instead of  $\omega$  for simplicity. Notice that this generalized momentum ( $\pi$ ) is consistent with the usual quantum mechanical procedure  $p \rightarrow p - eA$  that is used in Yang–Mills theories, as well as with the usual GR expression  $p_\alpha = mg_{\alpha\beta}v^\beta$ . Now it is easy to recognize the mass-shell constraint as a mathematical identity:

$$\frac{\vec{v}}{\sqrt{\vec{v}^2}} \cdot \frac{\vec{v}}{\sqrt{\vec{v}^2}} = 1 \Rightarrow \pi_\alpha \pi^\alpha = m^2 \Rightarrow \left( \vec{p} - e \vec{A} - \vec{S}_3(v) - \vec{S}_4(v) - \dots \right)^2 \Psi = m^2 \Psi.$$

Notice that “gravity” as represented by the metric is gone, while the Klein–Gordon equation appears. The  $v$  dependence in the  $S$  terms reminds us about the problem related to the change of coordinates  $(x, v) \rightarrow (x, p)$ . So, at this stage we may proceed with the Klein–Gordon equation, if we wish.

**Dirac equation from  $H=0$ .** An interesting approach to the Dirac equation has been suggested by H. Rund.[4] The idea uses Hamiltonian linear in the momentum ( $H = \gamma^\alpha p_\alpha$ ) and the base manifold principle group  $G$ . To have the Hamiltonian  $H$  invariant under  $G$ -transformations, the  $\gamma$  objects should transform appropriately and provide also a realization of the generators of  $G$ . Since we want  $\gamma$  and  $p$  to transform as vectors, it is clear that  $p$  should be a covariant derivative, but what is its structure? Consider a homogeneous Lagrangian that can be written as  $L(\phi, \omega) = \omega^\Gamma p_\Gamma = \omega^\Gamma \partial L(\phi, \omega) / \partial \omega^\Gamma$  with a Hamiltonian function that is identically zero:  $h = \omega^\Gamma \partial L(\phi, \omega) / \partial \omega^\Gamma - L(\phi, \omega) \equiv 0$ . Notice that  $\omega^\Gamma$  is the determinant of a matrix (the Jacobian of a transformation[21]); thus  $\omega^\Gamma \rightarrow \gamma^\Gamma$  seems an interesting option for quantization. Even more, for the Dirac theory we know that  $\gamma^\alpha$  are the ‘velocities’ ( $dx/d\tau = \partial H / \partial p$ ).

If we quantize using ( $h \rightarrow H$ ), then the space of physical states should satisfy:  $H\Psi = 0$ . By applying  $\omega^\Gamma \rightarrow \gamma^\Gamma$ , which means that the (generalized) velocity is considered as a vector with non-commutative components, we have  $(\gamma^\Gamma p_\Gamma - L(\phi, \gamma))\Psi = 0$ . For a point particle, using the canonical form of the Lagrangian (4) and the algebra of the  $\gamma$  matrices following Run's approach[4] this gives:

$$\begin{aligned} H &= \gamma^\alpha p_\alpha - L(\phi, \gamma) = \gamma^\alpha p_\alpha - eA_\alpha \gamma^\alpha - m\sqrt{g_{\alpha\beta}\gamma^\alpha\gamma^\beta} - \dots \sqrt[m]{S_m(\vec{\gamma}, \dots, \vec{\gamma})}, \\ &\rightarrow \gamma^\alpha p_\alpha - eA_\alpha \gamma^\alpha - m - \dots \sqrt[2m]{S_{2m}g^m} - \dots \sqrt[2n+1]{S_{2n+1}g^n}\gamma\dots \end{aligned}$$

Since  $g_{\alpha\beta}$  is a symmetric tensor such that  $\{\gamma^\alpha, \gamma^\beta\} \sim g^{\alpha\beta}$ , then  $g_{\alpha\beta}\gamma^\alpha\gamma^\beta \sim g_{\alpha\beta}\{\gamma^\alpha, \gamma^\beta\} \sim g_{\alpha\beta}g^{\alpha\beta} \sim 1$ . Therefore, gravity seems to leave the picture again. The symmetric structure of the extra terms  $S_m$  can be used to reintroduce  $g$  and to reduce the powers of  $\gamma$ . Thus the higher even order terms contribute to the mass  $m$ , making it variable[22] with  $\vec{x}$ .

**Summary.** We have discussed the structure of the matter Lagrangian for extended objects. Imposing reparametrization invariance of the action  $S$  naturally leads to a first order homogeneous Lagrangian. In its canonical form,  $L$  contains electromagnetic and gravitational interactions, as well as interactions that are not yet identified. The non-relativistic limit for a brane has been defined as those coordinates where the brane is an integral sub-manifold of the target space. This gauge can be used to remove reparametrization invariance of the action  $S$  and make the Hamiltonian function suitable for canonical quantization. The existence of a mass-shell constraint is universal. It is essentially due to the gravitational (quadratic in velocities) type interaction in the Lagrangian and always leads to a Klein-Gordon like equation. Once the algebraic properties of the  $\gamma$ -matrices are defined, one can use  $v \rightarrow \gamma$  quantization in the Hamiltonian function  $h = pv - L(x, v)$  to obtain the Dirac equation.

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